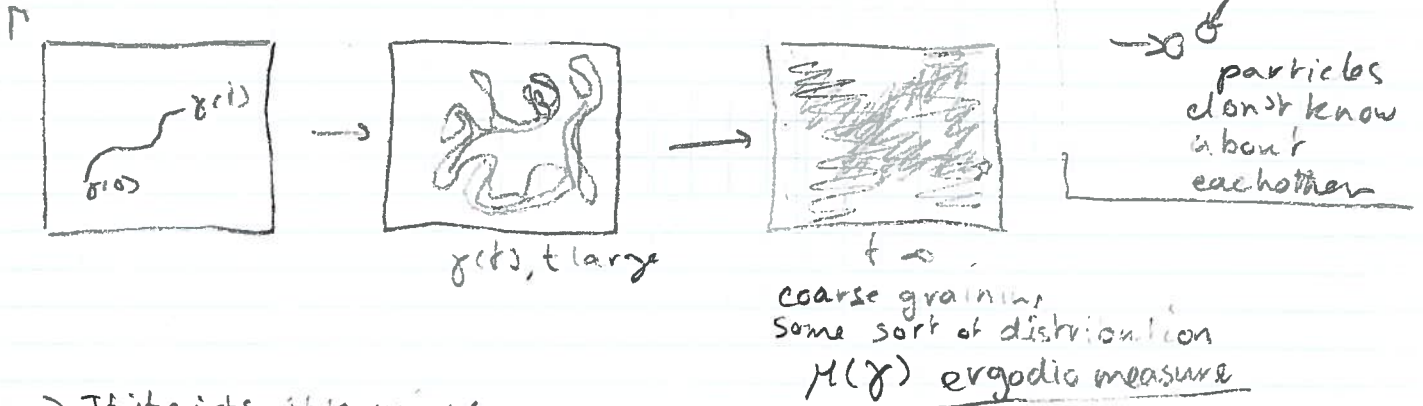


Stat Phys! averages

Boltzmann: reversible microscopic \leftrightarrow irreversible decay to equilibrium (21)
statistical assumptions,

Ergodicity Where does it go after 'long time?' coarse graining, Stoßzahlansatz



• \Rightarrow If it exists, it's unique

$\langle A \rangle_{\text{time}} = \langle A \rangle_{\mu(\gamma)}$ time average = ensemble average.

• hard to prove, ergodic/chaotic hypothesis

- Theoretical :
- equilibrium exists
 - $\mu(\gamma)$ tells us what it looks like (not everything)
 - \Rightarrow Equilibrium Stat Phys.
- Computational :
- allows for reduction of system (MC)
 - can induce various ensembles to represent physical situations
 - helps with calculation of averages

Watch out: not all systems are really ergodic on your time scales

INSERT AVERAGING HERE -
 Constant parameters

N V E T P μ

\hookrightarrow what we have been doing so far, microcanonical

• $\mu(\gamma)$ says: all states equally likely.
 equal amounts of time in equal amounts of phase space
 Ω # states with energy E , or volume of energy shell

entropy $S = k_B \log \Omega$

macroscopic quantities go to most common values, i.e. $\max S$

\Rightarrow second law of thermodynamics.

$T = \left(\frac{\partial S}{\partial E} \right)^{-1}_{N,V}$ $\mu = -T \left(\frac{\partial S}{\partial N} \right)_{E,V}$ $P = T \left(\frac{\partial S}{\partial V} \right)_{E,N}$

1st law: $dE = T dS - P dV + \mu dN$.

Determination of averages etc. also to do with correlations ^{error in}

often skatt expressed as averages

f.e. $C_V = \left. \frac{dE}{dT} \right|_{N,V}$

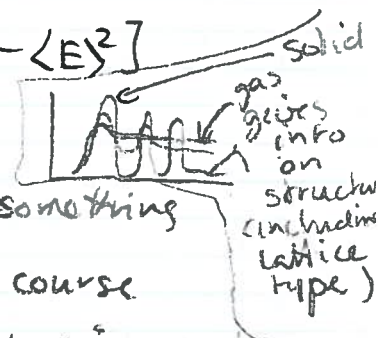
$Z = \frac{1}{N!} \sum \exp(-\beta E_i)$

$= -\beta^2 \frac{\partial}{\partial \beta} \langle E \rangle$

$\langle E \rangle = \frac{1}{N! Z} \sum E_i \exp(-\beta E_i)$

$= -\beta^2 \left\{ \frac{1}{N! Z} \sum E_i^2 \exp(-\beta E_i) + \frac{1}{N! Z^2} \left[\sum E_i \exp(-\beta E_i) \right]^2 \right\} = \beta^2 [\langle E^2 \rangle - \langle E \rangle^2]$

Similar, pressure $\langle p \rangle$ at boundary etc. --
RDF is average populations etc.



Some things are harder (not derivatives of 2 or something else obvious)

FOR G, S, μ \rightarrow maybe more later on in course

Evaluating averages is easy. Error due to finite size

finite time

$\langle A \rangle$ is average of A, SD $\sigma = \sqrt{\langle A^2 \rangle - \langle A \rangle^2}$ \leftarrow finite time

time averages, temporal correlation, so not independent realisations.

(so not just error = $\sigma / \sqrt{\# \text{ terms averaged}}$)

estimate the correlation time via (or length) correlation-function

\leftarrow should be # independent samples

$C_{AA}(k) = \langle (A_n - \langle A \rangle)(A_{n+k} - \langle A \rangle) \rangle$ $C_{AA}(0) = \sigma^2$, then falls off

number of ways to get corr time τ from this.

- fit to exp part,

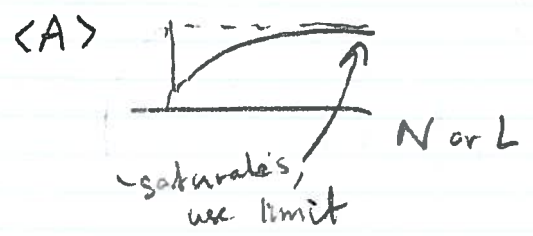
$-\tau = \frac{1}{2} \sum_{n=-\infty}^{\infty} \frac{C_{AA}(n)}{C_{AA}(0)}$ (same for C_{AA} exponential in n)

estimating error (see book) \Rightarrow error² = $\frac{2\tau}{\# \text{ samples}} \sigma^2$

Alternatively: just divide into large blocks of length $> \tau$, calculate separate averages for each.

Now you have small number of independent samples.

finite size systematic error



(assuming for $N \rightarrow \infty$ not dependent on N)

hard to keep E constant in experiment (heat bath)

T instead \Rightarrow canonical ensemble. NVT constant.
most common

$\mu(\gamma) \propto \exp(-\beta E)$ Boltzmann factor. $\beta = \frac{1}{T} = \frac{1}{k_B T}$

\Rightarrow normalisation with partition function

$\langle A \rangle_{NVT} = \frac{1}{N! Z} \sum_{\text{states}} A e^{-\beta E(\text{state})}$

$Z(N, V, T) = \frac{1}{N!} \sum_{\text{states}} e^{-\beta E(\text{state})} = \sum_E e^{-\beta E} \Omega(N, V, E) = \sum_E e^{-\beta(E-TS)}$

$F = T \log Z = E - TS.$

minimised in equilibrium in canonical ensemble

There are 2 more

- isothermal, isobaric Gibbs. $G = E - TS + PV$
- grand canonical

NPT pressure constant $\mu(\gamma) = \exp(-\beta E - \beta PV)$
 μVT open system $(-\beta E - \beta \mu N)$

How to achieve NVT in simulations (or part of simulated system): must couple to effective heat bath. Thermostats

- experiments / real systems have heat baths.
- sometimes you need to remove energy, because you are putting it in
- non-physical constructs: take extreme care
- examples: Langevin, Nosé-Hoover

end lecture 6

Langevin - dampings. random force gaussian uncorrelated noise. $\dot{v} = -\eta v + a(t) + \xi(t)$. $F = \eta m v = \gamma v$

NOT DETERMINISTIC

drift a_2

assumptions: bath is fast & correlation decays exponentially
adiabatic, dampings is viscous.

example: 1D system, $V(x)$



Fokker-Planck equation $p(x, v)$ probability density

$\frac{d}{dt} p(x, v) = \frac{\partial}{\partial x} [a(x) - \eta v] p(x, v) + \frac{\partial}{\partial v} [v p(x, v)] + \frac{\partial^2}{\partial v^2} p(x, v)$
drift a_1 drift a_2 diffusion

drift: $\frac{dp(v)}{dv} = \frac{d}{dt} [p(v+dv) - v p(v)]$