

Stat Phys! averages

Boltzmann:

reversible microscopic  $\leftrightarrow$  irreversible decay to equilibrium  
statistical assumptions

(10)

Ergodicity

Where does it go after long time?

coarse grainings, Stoßzahlansatz



particles don't know about each other

coarse grainings, some sort of distribution  $\mu(\gamma)$  ergodic measure

$\Rightarrow$  If it exists, it's unique

$\langle A \rangle_{\text{time}} = \langle A \rangle_{\mu(\gamma)}$  time average = ensemble average.

hard to prove, ergodic/chaotic hypothesis

Theoretical

- equilibrium exists
- $\mu(\gamma)$  tells us what it looks like (not everything)
- $\Rightarrow$  Equilibrium Stat Phys.

Computational

- allows for reduction of system (MC)
- can induce various ensembles to represent physical situations
- helps with calculation of averages

Watch out: not all systems are really ergodic on your time scales

~~INSERT AVERAGING HERE (next page)~~  
 Constant parameters

N V E T P  $\mu$

what we have been doing so far, microcanonical

$\mu(\gamma)$  says: all states equally likely.

equal amounts of time in equal amounts of phase space  
 $\Omega$  # states with energy  $E$ , or volume of energy shell

entropy  $S = k_B \log \Omega$

macroscopic quantities go to most common values, i.e. max  $S$

$\Rightarrow$  second law of thermodynamics.

$T = \left( \frac{\partial S}{\partial E} \right)_{N,V}^{-1}$      $\mu = -T \left( \frac{\partial S}{\partial N} \right)_{E,V}$      $P = T \left( \frac{\partial S}{\partial v} \right)_{E,N}$

1st law:  $dE = T dS - P dV + \mu dN$

Determination of averages etc. also to do with correlations <sup>error in</sup>

often still expressed as averages

f.e.  $C_V = \left. \frac{dE}{dT} \right|_{N,V}$

$= -\beta^2 \frac{d}{d\beta} \langle E \rangle$

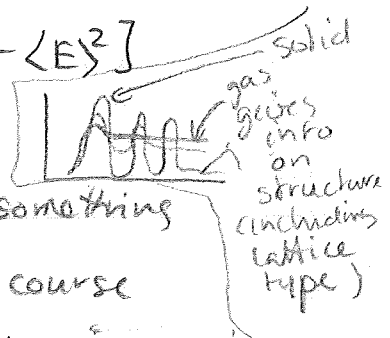
$= -\beta^2 \left\{ \frac{1}{N! 2} \sum_i E_i^2 \exp(-\beta E_i) \right.$

$\left. + \frac{1}{N! 2^2} \left[ \sum_i E_i \exp(-\beta E_i) \right]^2 \right\} = \beta^2 [\langle E^2 \rangle - \langle E \rangle^2]$

$Z = \frac{1}{N!} \sum_i \exp(-\beta E_i)$

$\langle E \rangle = \frac{1}{N! Z} \sum_i E_i \exp(-\beta E_i)$

Similar, pressure  $\langle p \rangle$  at boundary etc. →



RDF is average populations etc. some things are harder (not derivatives of 2 or something else obvious)

FOR G, S,  $\mu$  → maybe more later on in course

evaluating averages is easy. Error due to finite size

finite time

$\langle A \rangle$  is average of A, SD  $\sigma = \sqrt{\langle A^2 \rangle - \langle A \rangle^2}$  ← finite time

time averages, temporal correlation, so not independent realisations.

(so not just error =  $\sigma / \sqrt{\# \text{ terms averaged}}$ )

↪ should be  $\#$  independent samples

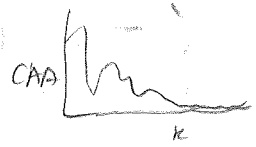
estimate the correlation time via (or length) correlation-function

$C_{AA}(k) = \langle (A_n - \langle A \rangle)(A_{n+k} - \langle A \rangle) \rangle$   $C_{AA}(0) = \sigma^2$ , then falls off

number of ways to get correlation  $\tau$  from this.

- fit exponential function

$-\tau = \frac{1}{2} \sum_{n=0}^{\infty} \frac{C_{AA}(n)}{C_{AA}(0)}$  (same for  $C_{AA}$  exponential in  $n$ )



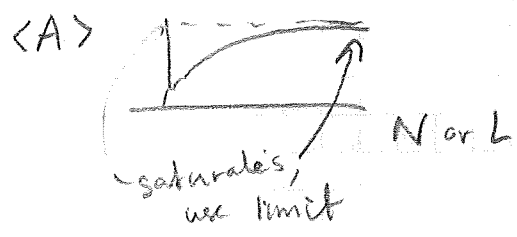
estimating error (see book)  $\Rightarrow$  error<sup>2</sup> =  $\frac{2\tau}{\# \text{ samples}} \sigma^2$

Alternatively: just divide into large blocks of length  $> \tau$ , calculate separate averages for each.

Now you have small number of independent samples.

block averages

finite size systematic error



(assuming for  $N \rightarrow \infty$  not dependent on  $N$ )

hard to keep E constant in experiment (heat bath)

T instead  $\Rightarrow$  canonical ensemble. most common NVT constant.

$\mu(x) \propto \exp(-\beta E)$   
Boltzmann factor,  $\beta = \frac{1}{T} = \frac{1}{k_B T}$

$\Rightarrow$  normalisation with partition function

$\langle A \rangle_{NVT} = \frac{1}{N! Z} \sum_{states} A e^{-\beta E(state)}$

$Z(N, V, T) = \frac{1}{N!} \sum_{states} e^{-\beta E(state)} = \sum_E e^{-\beta E} \Omega(N, V, E) = \sum_E e^{-\beta(E-TS)}$

$F = T \log Z = E - TS.$

minimised in equilibrium in canonical ensemble

There are 2 more

- isothermal, isobaric Gibbs,  $G = E - TS + PV$
- grand canonical

NPT pressure constant  
 $\mu(x) = \exp(-\beta E - \beta PV)$   
 NVT open system  $(-\beta E - \beta \mu N)$

end below 6

How to achieve NVT in simulations (or part of simulated system): must couple to effective heat bath. Thermostats

- experiments / real systems have heat baths.
- sometimes you need to remove energy, because you are putting it in
- non-physical constructs: take extreme care
- examples: Langevin, Nosé-Hoover

Langevin dampens random force gaussian uncorrelated noise  
 $\dot{v} = -\eta v + a(t) + \xi(t)$   
 $F = \eta m v = \gamma v$

NOT DETERMINISTIC

drift  $a_2$

assumptions: bath is fast & correlation decays exponentially  
adiabatic, damping is viscous.

example: 1D system,  $V(x)$



Fokker-Planck equation  $p(x, v)$  probability density

$\frac{d}{dt} p(x, v) = \frac{d}{dx} [a(x) - \eta v] p(x, v) + \frac{\partial}{\partial v} \frac{\partial}{\partial v} D_{vv} p(x, v)$   
 drift out, drift in, diffusion

drift:  $\frac{dp(x)}{dx} = \frac{d}{dt} [p(v+dv) - v p(x)]$